

脈動オーロラにおける内部変調の起源について

三好 由純 [1]; 齊藤 慎司 [2]; 関 華奈子 [1]; 西山 尚典 [3]; 片岡 龍峰 [3]; 浅村 和史 [4]; 加藤 雄人 [5]; 海老原 祐輔 [6]; 坂野井 健 [7]; 平原 聖文 [8]; 大山 伸一郎 [9]; 栗田 怜 [10]; 尾崎 光紀 [11]; Santolik Ondrej [12]
[1] 名大 STE 研; [2] 名大理; [3] 極地研; [4] 宇宙研; [5] 東北大・理・地球物理; [6] 京大生存圏; [7] 東北大・理; [8] 名大・STE 研; [9] 名大・太陽研; [10] 名古屋大・STEL; [11] 金沢大・理工・電情; [12] チェコ科学アカデミー

Origin of internal modulations of the pulsating aurora

Yoshizumi Miyoshi[1]; Shinji Saito[2]; Kanako Seki[1]; Takanori Nishiyama[3]; Ryuho Kataoka[3]; Kazushi Asamura[4]; Yuto Katoh[5]; Yusuke Ebihara[6]; Takeshi Sakanoi[7]; Masafumi Hirahara[8]; Shin-ichiro Oyama[9]; Satoshi Kurita[10]; Mitsunori Ozaki[11]; Ondrej Santolik[12]
[1] STEL, Nagoya Univ.; [2] Nagoya Univ.; [3] NIPR; [4] ISAS/JAXA; [5] Dept. Geophys., Grad. Sch. Sci., Tohoku Univ.; [6] RISH, Kyoto Univ.; [7] Grad. School of Science, Tohoku Univ.; [8] STEL, Nagoya Univ.; [9] STEL, Nagoya Univ.; [10] STEL, Nagoya Univ.; [11] Electrical and Computer Eng., Kanazawa Univ.; [12] The Czech Academy of Sciences

We investigate the origin of the internal modulations, so-called quasi-few Hz modulations of precipitating electrons for the pulsating aurora by comparative studies between the Reimei observation and a computer simulation for the chorus-wave particle interactions. From the Reimei observations, the main modulation of precipitation, with a few seconds, and the internal modulations, with a few Hz, that are embedded inside the main modulations are identified. A computer simulation that calculates interactions between the bouncing electrons along the magnetic field and the propagating whistler mode waves shows that the generation and collapse of the lower-band chorus bursts determines on-off switching of the pulsating aurora. A train of rising tone elements embedded in the lower-band chorus bursts drives the internal modulations. Besides the internal modulations, the Reimei satellite found a precipitation gap between intermittent precipitations above a few keV and stable precipitations around 1 keV. The gap corresponds to the half-gyro frequency gap between the lower-band chorus waves and the upper-band chorus waves.