

Density depletions associated with enhancements of ECH emissions observed by ERG

Yoichi Kazama[1]; Hirotsugu Kojima[2]; Yoshizumi Miyoshi[3]; Yoshiya Kasahara[4]; Hideyuki Usui[5]; B.-J. Wang[6]; S.-Y. Wang[7]; Sunny W. Y. Tam[8]; Tzu-Fang Chang[9]; Paul Ho[10]; Kazushi Asamura[11]; Atsushi Kumamoto[12]; Fuminori Tsuchiya[13]; Yasumasa Kasaba[14]; Shoya Matsuda[15]; Masafumi Shoji[3]; Ayako Matsuoka[16]; Mariko Teramoto[17]; Takeshi Takashima[18]; Iku Shinohara[19]

[1] ASIAA; [2] RISH, Kyoto Univ.; [3] ISEE, Nagoya Univ.; [4] Kanazawa Univ.; [5] System informatics, Kobe Univ.; [6] ASIAA, Taiwan; [7] ASIAA, Taiwan; [8] ISAPS, NCKU, Taiwan; [9] ISEE, Nagoya Univ.; [10] Institute of Astronomy and Astrophysics, Academia Sinica, Taiwan; [11] ISAS/JAXA; [12] Dept. Geophys, Tohoku Univ.; [13] Planet. Plasma Atmos. Res. Cent., Tohoku Univ.; [14] Tohoku Univ.; [15] ISAS/JAXA; [16] ISAS/JAXA; [17] ISEE, Nagoya University; [18] ISAS, JAXA; [19] ISAS/JAXA

Small-scale density depletions of cold electrons associated with electrostatic electron cyclotron harmonic (ECH) waves were observed by the ERG spacecraft during a plasmopause crossing near the magnetic equator in the post-midnight. During this event, a hot electron component (several tens eV to ~ 1 keV) and an energetic electron component (> 1 keV) were measured, and both of the electron components showed pancake-like velocity distributions. The total electron density derived from the local upper-hybrid resonance (UHR) frequency showed roughly two orders of magnitude larger than that of the hot electrons, indicating existence of a cold and dense electron population below ~ 20 eV. The cold electron density variation was well anti-correlated to the intensity of the ECH emissions, which means that the ECH emissions were intensified inside a density depletion region (DDR) of the cold and dense electrons. Moreover, a flux enhancement of hot electrons in the perpendicular direction was also associated with the ECH emission intensity. Lower-energy hot electrons (e.g., ~ 100 eV) show a better correlation with the ECH emission intensity compared with higher-energy electrons (e.g., ~ 1 keV). The relation between the perpendicular electrons and the ECH emission intensity suggests energization of electrons in the perpendicular directions by electric field oscillation of the ECH waves.