

S001-05

A 会場 : 11/4 PM1 (13:45-15:30)

14:55~15:10

高強度レーザー駆動多種イオンプラズマ中の無衝突静電衝撃波によるイオン加速

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High-intensity laser driven ion acceleration by collisionless electrostatic shock in a multicomponent ion plasma

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Shock waves observed in astrophysical systems, for example, supernova remnant shock and the bow shock of the earth, are shocks in collisionless plasmas. Such collisionless shocks are the most promising candidate for the generation mechanism of cosmic-rays. The interaction between the electric and magnetic fields, generated by the collective phenomenon of plasma, and charged particles is important for the generation of collisionless shocks. By using a high-intensity laser, an acceleration mechanism called ion acceleration by a laser-driven collisionless electrostatic shock is drawing attention [1-4]. In the collisionless electrostatic shock ion-acceleration (CESA), upstream ions of the shock are reflected by the shock potential and accelerated.

The CESA experiments are performed with high-intensity LFEX laser beams (pulse width = 1.5 ps, energy ~ 300 J, laser intensity $\sim (3-6) \times 10^{19}$ W/cm², and normalized laser intensity $a_0 \sim 2$) at the Institute of Laser Engineering, Osaka University. In order to generate an initial plasma with a near-critical density and a long scale-length on the rear-side of the drive-laser irradiated target, which is suitable for CESA [2-4], the rear surface of the target is irradiated with an ionization laser at 2.5 ps before the drive laser. As the ionization laser, one of the Gekko XII laser beams (pulse width = 1.3 ps, energy ~ 3 J, laser intensity $\sim 3 \times 10^{11}$ W/cm²) is focused on the back surface of the target (a thin foil of C₈H₇Cl with a thickness of 1 μ m). Thomson parabola spectrometer (TPS) and electron spectrometer (ESM), located on the rear-side of the target, are used to measure ion and electron spectrum, respectively. We observed clear evidences for the proton acceleration by the collisionless electrostatic shock.

[1] D. Haberberger, et al, Nature Phys. **8**, 95 (2012).

[2] R. Kumar, Y. Sakawa, et al, Phys. Rev. Accel. Beams **22**, 043401 (2019).

[3] R. Kumar, Y. Sakawa, et al, Phys. Rev. E **103**, 043201 (2021).

[4] Y. Sakawa, Y. Ohira, et al, Phys. Rev. E **104**, 055202 (2021).