## れいめい衛星による低緯度大気光ならびにスプライト観測

# 坂野井 健 [1]; 山崎 敦 [2]; 藤原 均 [3]; 佐藤 光輝 [4]; 高橋 幸弘 [3]; 岡野 章一 [5]; 平原 聖文 [6] [1] 東北大・理; [2] 宇宙科学研究本部; [3] 東北大・理・地球物理; [4] 北大; [5] 東北大・理; [6] 東大・理・地惑

## Airglow and sprites observations at low-latitudes by Reimei

# Takeshi Sakanoi[1]; Atsushi Yamazaki[2]; Hitoshi Fujiwara[3]; Mitsuteru Sato[4]; Yukihiro Takahashi[3]; Shoichi Okano[5]; Masafumi Hirahara[6]

[1] PPARC, Grad. School of Sci., Tohoku Univ.; [2] ISAS/JAXA; [3] Dept. of Geophysics, Tohoku Univ.; [4] Hokkaido Univ.;
[5] PPARC, Tohoku Univ.; [6] Dept. Earth & Planet. Sci, Univ. Tokyo

http://pparc.geophys.tohoku.ac.jp/~tsakanoi/

It is important to observe airglow emission since airglow emission represents the dynamics and chemistry in the thermosphere. So far, the measurements of airglow were mainly carried out from ground-based instruments. However, there are some difficulties in past measurements: the field-of view is limited within hundreds kilo-meters and the optical measurements depend on the weather conditions. On the other hand, a satellite observation enables us to perform the continuous measurement of airglow distribution in the wide-range. However, satellite data are rarely found for visible airglow emissions.

The Reimei satellite carries a multi-spectral auroral camera (MAC) which can measure 428 nm (N2+ 1NG), 558 nm (OI) and 670 nm (OH Meinel band / N2 1PG) emissions. From March 2008, the observation of airglow and sprite emissions started with MAC by pointing the field-of-view toward the earth's limb at middle and low-latitudes. Since then, the observation continues routinely every two or three times a week. Simultaneous observations between 428 nm and 670 nm or 558 nm and 670 nm are performed with a time resolution of 1 sec and a spatial resolution at a tangential point of  $^{4}$  km, respectively. According to the noon-midnight sun-synchronous orbit of Reimei at an altitude of  $^{640}$  km, the observation is made around the midnight sector.

From the preliminary analysis of MAC data obtained at ~1913 UT on Apr. 15, 2008, it is found that the vertical profile of O 558 nm shows the maximum (~4kR) at an altitude of 95 km, and that the second peak (0.3 kR) is also seen at an altitude of ~220 km suggesting the existence of different emission process in the F-region. The peak of OH 670 nm emission intensity is ~2kR. Considering that the filter bandwidth of 670 nm is 38 nm, the OH band intensity is estimated to be 53 R/nm on average. We will estimate the volume emission rate by using these observed vertical profiles. In addition, we found the sprite event at 1915:49.19 UT in this path. It is expected that electron enery of sprites is estimated using simultaneous imaging data at N2+ 1NG 428nm and at N2 1PG 670 nm. In this presentation, recent results of airglow and sprites observations at low-latitudes by Reimei will be presented.