Multi-wavelength observations of Jovian aurora coordinated with Hisaki and other space telescopes

Tomoki Kimura[1]; Sarah Badman[2]; Chihiro Tao[3]; Kazuo Yoshioka[4]; Go Murakami[2]; Atsushi Yamazaki[5]; Fuminori Tsuchiya[6]; Ralph Kraft[7]; Graziella Branduardi-Raymont[8]; Yuichiro Ezoe[9]; Masaki Fujimoto[10]

 [1] JAXA/ISAS; [2] ISAS/JAXA; [3] LPP, Ecole Polytechnique; [4] JAXA/ISAS; [5] ISAS/JAXA; [6] Planet. Plasma Atmos. Res. Cent., Tohoku Univ.; [7] Harvard Smithonian Astronomy Observatory; [8] University Collage London; [9] Tokyo Metropolitan University; [10] ISAS, JAXA

From January to April 2014, two observing campaigns by multi-wavelength remote sensing from X-ray to radio were performed to uncover energy transport process in Jupiter's plasma environment using space telescopes and ground-based facilities. These campaigns were triggered by the new Hisaki spacecraft launched in September 2013, which is an extremely ultraviolet (EUV) space telescope of JAXA designed for planetary observations.

In the first campaign in January, Hubble Space Telescope (HST) made imaging of far ultraviolet (FUV) aurora with a high special resolution (0.08 arcsec) through two weeks while Hisaki continuously monitored aurora and plasma torus emissions in EUV wavelength with a high temporal resolution (more than 1 min). We discovered new magnetospheric activities from the campaign data: e.g., internally-driven type auroral brightening associated with hot plasma injection, and plasma and electromagnetic filed modulations in the inner magnetosphere externally driven by the solar wind modulation.

The second campaign in April was performed by Chandra X-ray Observatory (CXO), XMM newton, and Suzaku satellite simultaneously with Hisaki. Relativistic auroral accelerations in the polar region and hot plasma in the inner magnetosphere were captured by the X-ray space telescopes simultaneously with EUV monitoring of aurora and plasma torus. Auroral intensity in EUV indicated a clear periodicity of 45 minutes whereas the periodicity was not evident in X-ray intensity although previous observations by CXO indicated clear 40-minute periodicity in the polar cap X-ray aurora.

In this presentation, we show remarkable scientific results obtained these campaigns.