## R006-25 A 会場 :11/7 AM1(9:00-10:30) 09:45~10:00

#Kumar Sandeep<sup>1)</sup>, 三好 由純<sup>1)</sup>, ジョルダノヴァ ヴァニア<sup>2)</sup>,Kistler Lynn M.<sup>1,3)</sup>,Porunakatu Radhakrishna Shreedevi<sup>1)</sup>, 浅村 和史<sup>4)</sup>, 横田 勝一郎<sup>5)</sup>, 笠原 惹<sup>6)</sup>, 風間 洋一<sup>7)</sup>,Wang S.-Y.<sup>7)</sup>,Tam Sunny W. Y.<sup>8)</sup>, 三谷 烈 史<sup>4)</sup>, 東尾 奈々<sup>4)</sup>, 桂華 邦裕<sup>6)</sup>,Park Inchun<sup>1)</sup>, 堀 智昭<sup>1)</sup>, 田 采祐<sup>1)</sup>, 松岡 彩子<sup>9)</sup>, 今城 峻<sup>9)</sup>, 篠原 育<sup>4)</sup> (<sup>1</sup>ISEE,NU, (<sup>2</sup>LANL, (<sup>3</sup>University of New Hampshire, (<sup>4</sup> 宇宙研/宇宙機構, (<sup>5</sup>Osaka University, (<sup>6</sup>University of Tokyo, (<sup>7</sup>ASIAA, Taiwan, (<sup>8</sup>In of Space and Plasma Science, National Cheng Kung University, (<sup>9</sup>京大・地磁気センター

## Plasma pressure distribution of ions and electrons in the inner magnetosphere during CIR and CME storms observed by Arase

#Sandeep Kumar<sup>1</sup>), Yoshizumi Miyoshi<sup>1</sup>), Vania Jordanova<sup>2</sup>), Lynn M. Kistler<sup>1,3</sup>), Shreedevi Porunakatu Radhakrishna<sup>1</sup>), Kazushi Asamura<sup>4</sup>), Shoichiro Yokota<sup>5</sup>), Satoshi Kasahara<sup>6</sup>), Yoichi Kazama<sup>7</sup>), S.-Y. Wang<sup>7</sup>), Sunny W. Y. Tam<sup>8</sup>), Takefumi Mitani<sup>4</sup>), Nana Higashio<sup>4</sup>), Kunihiro Keika<sup>6</sup>), Inchun Park<sup>1</sup>), Tomoaki Hori<sup>1</sup>), ChaeWoo Jun<sup>1</sup>), Ayako Matsuoka<sup>9</sup>), Shun Imajo<sup>9</sup>), Iku Shinohara<sup>4</sup>)

<sup>(1</sup>ISEE,Nagoya Univ.,<sup>(2</sup>Space Science and Application, LANL,<sup>(3</sup>ISEOS, University of New Hampshire,<sup>(4</sup>ISAS/JAXA,<sup>(5</sup>Osaka Univ.,<sup>(6</sup>The University of Tokyo,<sup>(7</sup>ASIAA,Taiwan,<sup>(8</sup>Institute of Space and Plasma Sciences, National Cheng Kung University, Taiwan,<sup>(9</sup>WDC for Geomagnetism, Kyoto, Kyoto University

Geomagnetic storms are the main component of space weather, and the main phase of the geomagnetic storms are driven by Coronal Mass Ejections (CMEs) or Corotating Interaction Regions (CIRs). It is well known that CME-driven storms and CIR-driven storms have different evolutions of the Sym-H and the ion distributions in the inner magnetosphere [Miyoshi and Kataoka, 2005]. Enhancement of the ring current is a typical feature of the geomagnetic storm and a global decrease in the H component of the geomagnetic field is observed during the main phase of the geomagnetic storm. The ring current represents a diamagnetic current driven by the plasma pressure in the inner magnetosphere. The plasma pressure is mainly contributed by protons in an energy range of a few to a few hundreds of keV. The O+ contribution is also important, and sometimes dominates H+ during the geomagnetic field during the November 2017 CIR-driven storm by comparing Ring current Atmosphere interactions Model with Self Consistent magnetic field (RAM-SCB) simulation, Arase in-situ plasma/particle data, and ground-based magnetometer data [Kumar et al., 2021]. It has been shown that ion and electron distributions of CME/CIR-driven storms are different, especially for recovery phase [Miyoshi and Kataoka, 2005]. In this study, we examine statistically the spatial and temporal distribution of electrons and ions pressure with different energies and their contribution to the depression of the magnetic field during main phase, early recovery and late recovery phase for selected CIR and CME storms using in situ plasma/particle data obtained by Arase.